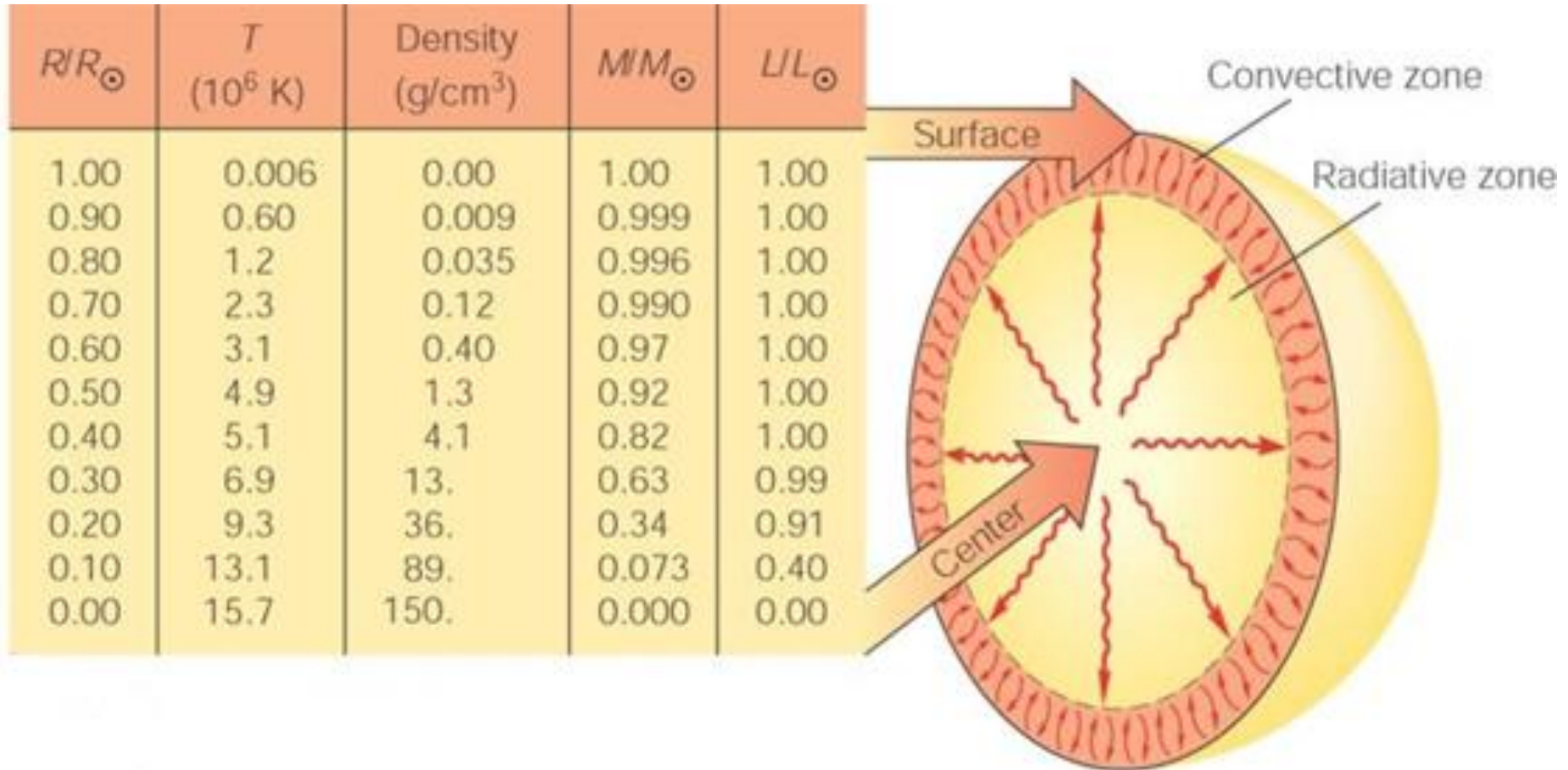


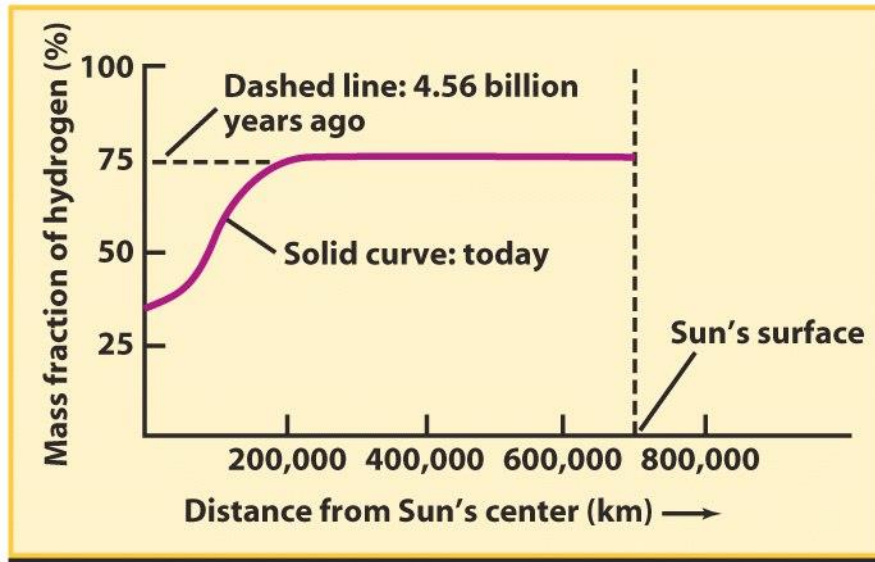
# The Internal Structure of the Sun



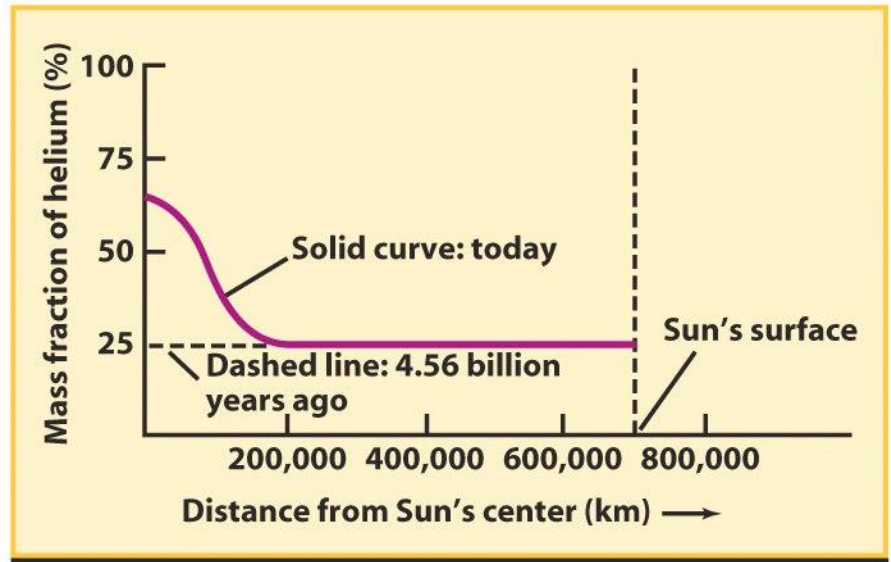
# The Sun: 4.5 billion years old

- The Sun has been a main-sequence star for 4.5 billion years, and at the core
  1. Hydrogen depleted by about 35%
  2. Helium amount increased
  3. Become 40% more luminous
  4. Has grown in radius by 6%
- The Sun should remain on the main sequence for another 7 billion years
- The Sun or  $1 M_{\odot}$  star has a main sequence lifetime of 12 billion years

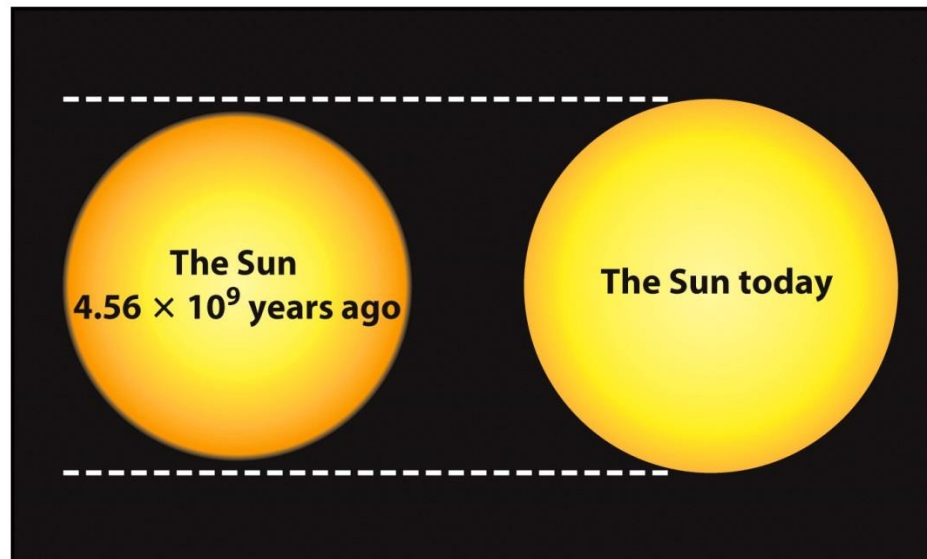
# The Sun: 4.5 billion years old



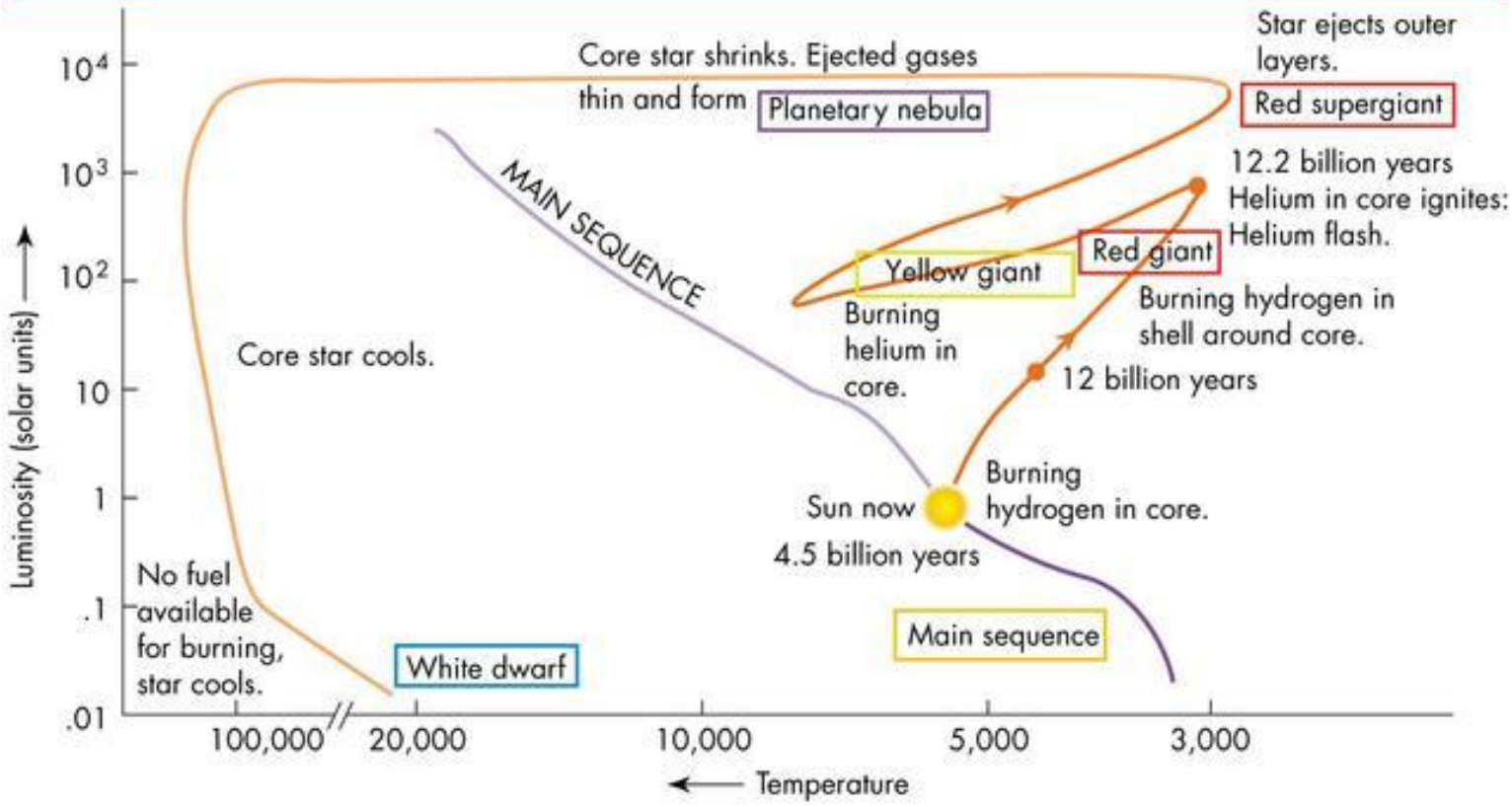
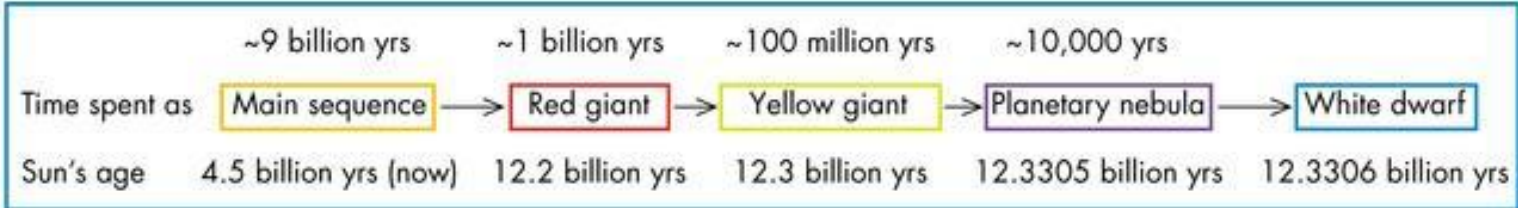
(a) Hydrogen in the Sun's interior



(b) Helium in the Sun's interior



# The Evolution of the Sun





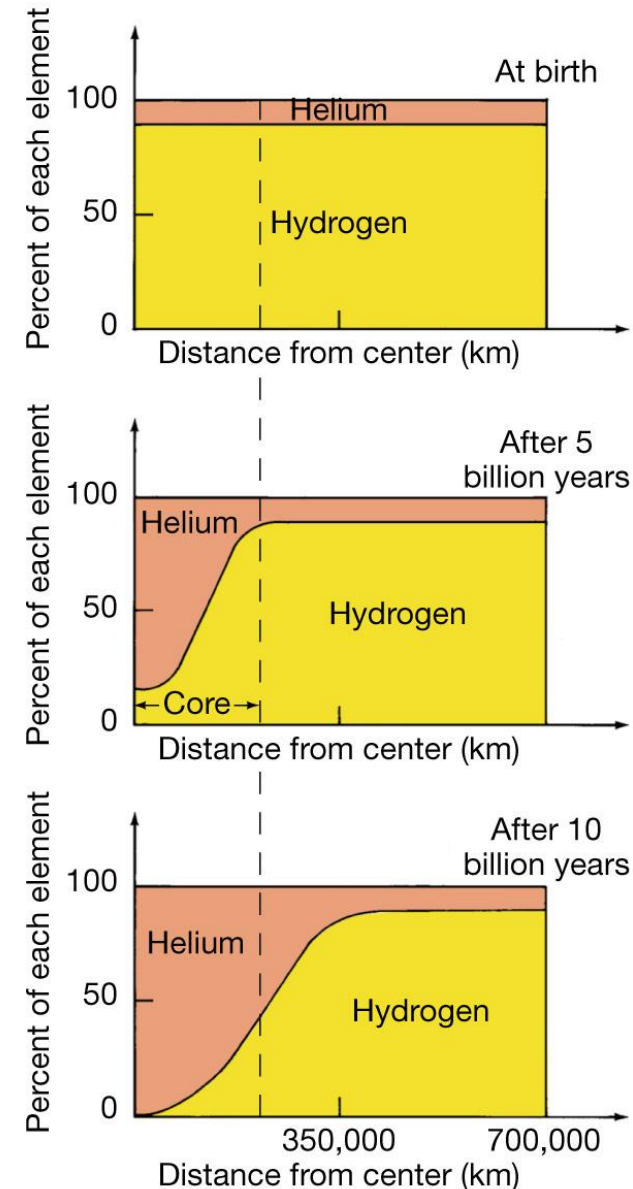
# Change of chemical composition

Asplund et al., 2009, Annual Review of Astronomy & Astrophysics, 47, 481

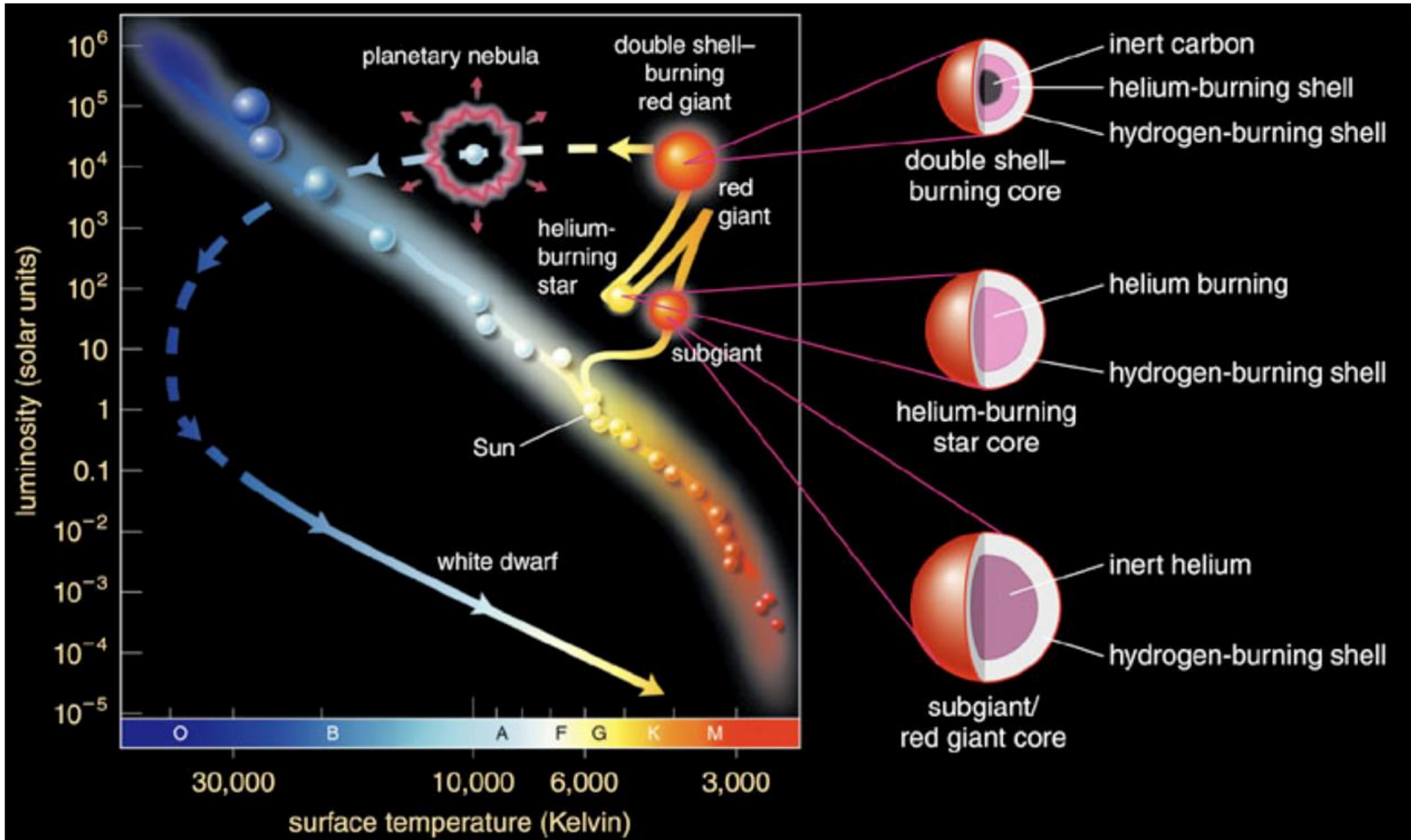
Table 4: The mass fractions of hydrogen (X), helium (Y) and metals (Z) for a number of widely-used compilations of the solar chemical composition.

Source	X	Y	Z	Z/X
<b>Present-day photosphere:</b>				
Anders & Grevesse (1989) <sup>a</sup>	0.7314	0.2485	0.0201	0.0274
Grevesse & Noels (1993) <sup>a</sup>	0.7336	0.2485	0.0179	0.0244
Grevesse & Sauval (1998)	0.7345	0.2485	0.0169	0.0231
Lodders (2003)	0.7491	0.2377	0.0133	0.0177
Asplund, Grevesse & Sauval (2005)	0.7392	0.2485	0.0122	0.0165
Lodders, Palme & Gail (2009)	0.7390	0.2469	0.0141	0.0191
Present work	0.7381	0.2485	0.0134	0.0181
<b>Proto-solar:</b>				
Anders & Grevesse (1989)	0.7096	0.2691	0.0213	0.0301
Grevesse & Noels (1993)	0.7112	0.2697	0.0190	0.0268
Grevesse & Sauval (1998)	0.7120	0.2701	0.0180	0.0253
Lodders (2003)	0.7111	0.2741	0.0149	0.0210
Asplund, Grevesse & Sauval (2005)	0.7166	0.2704	0.0130	0.0181
Lodders, Palme & Gail (2009)	0.7112	0.2735	0.0153	0.0215
Present work	0.7154	0.2703	0.0142	0.0199

<sup>a</sup> The He abundances given in Anders & Grevesse (1989) and Grevesse & Noels (1993) have here been replaced with the current best estimate from helioseismology (Sect. 3.9).



# Next Phases



# Subgiant Phase

- ***Core hydrogen fusion ceases*** when the hydrogen has been exhausted in the core of a main-sequence star
- This leaves a ***core of nearly pure helium***
- The ***core shrinks*** under self-gravitation due to the loss of hydrostatic equilibrium
- The ***core*** becomes ***hotter***
- ***Shell hydrogen fusion*** occurs just outside the core
- ***No fusion*** in the ***helium core***
- ***Subgiant Phase (SGB)***

# Red Giant Phase

- ***Shell hydrogen fusion*** works its way outward in the star and ***adds more helium*** into the ***core***
- ***Core becomes hotter***
- Shell hydrogen ***fusion*** occurs at a ***greater rate***
- ***Outer layers expands*** because of the increased energy flow
- At some point the core reaches the ***Schönberg-Chandrasekhar limit***

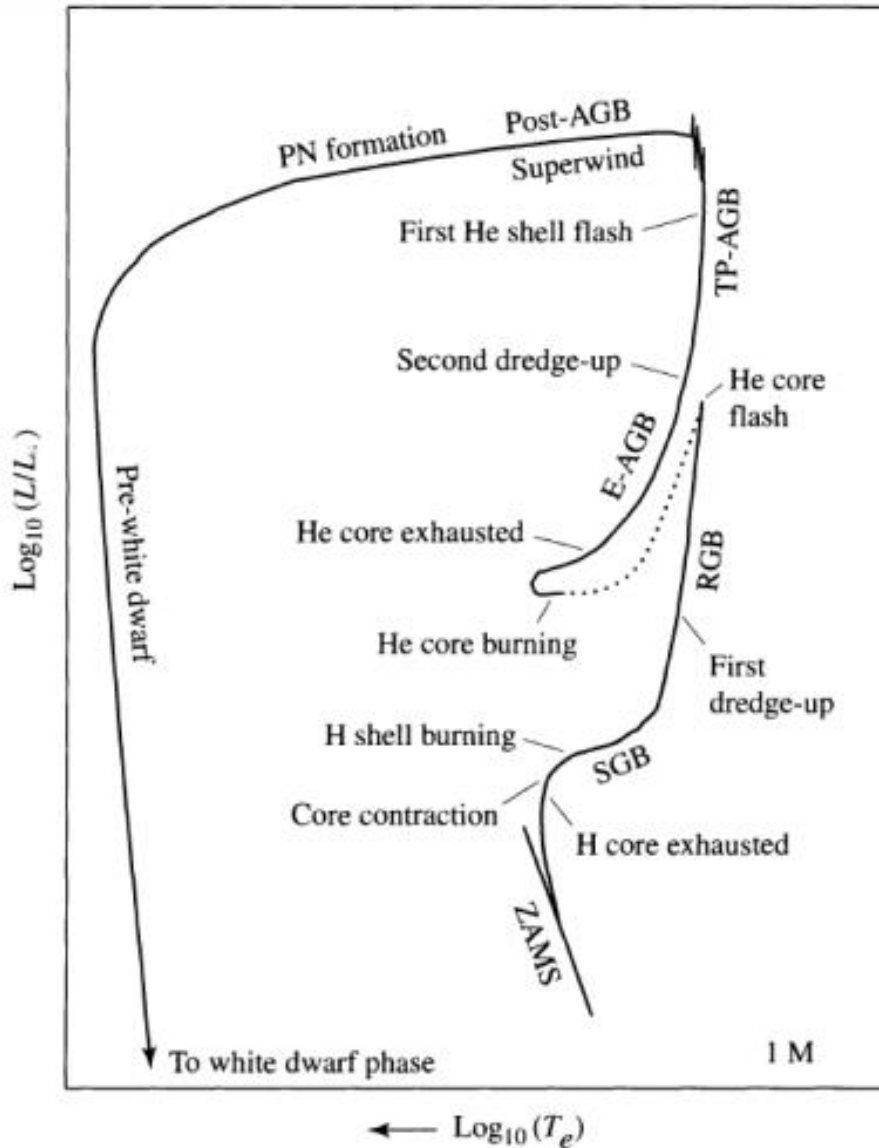


# Schönberg-Chandrasekhar limit

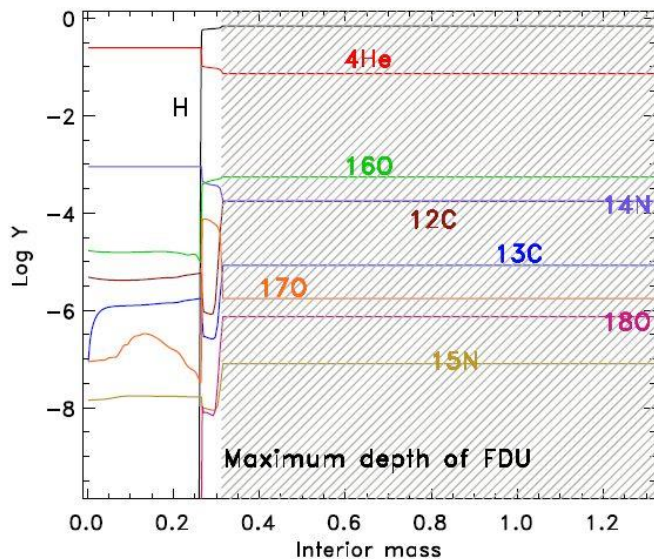
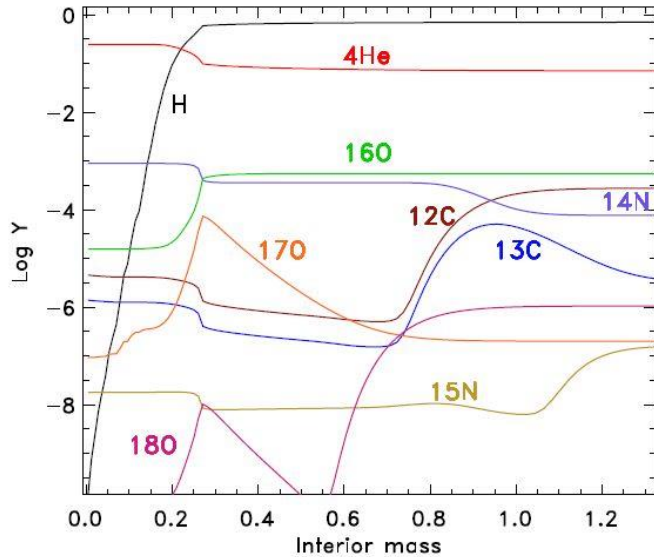
- The ***mass limit*** for the ***core*** which is capable of ***supporting*** the ***gravitational pressure*** of ***core and envelope***
- Core too massive => rapid contraction
- Release of a lot of gravitational potential energy
- Dumped into the envelope => heating up
- The Subgiant is now becoming a ***Red Giant (Red Giant Branch, RGB)***

# First dredge-up

As the outer atmosphere expands it becomes *convective* => *Mixing*



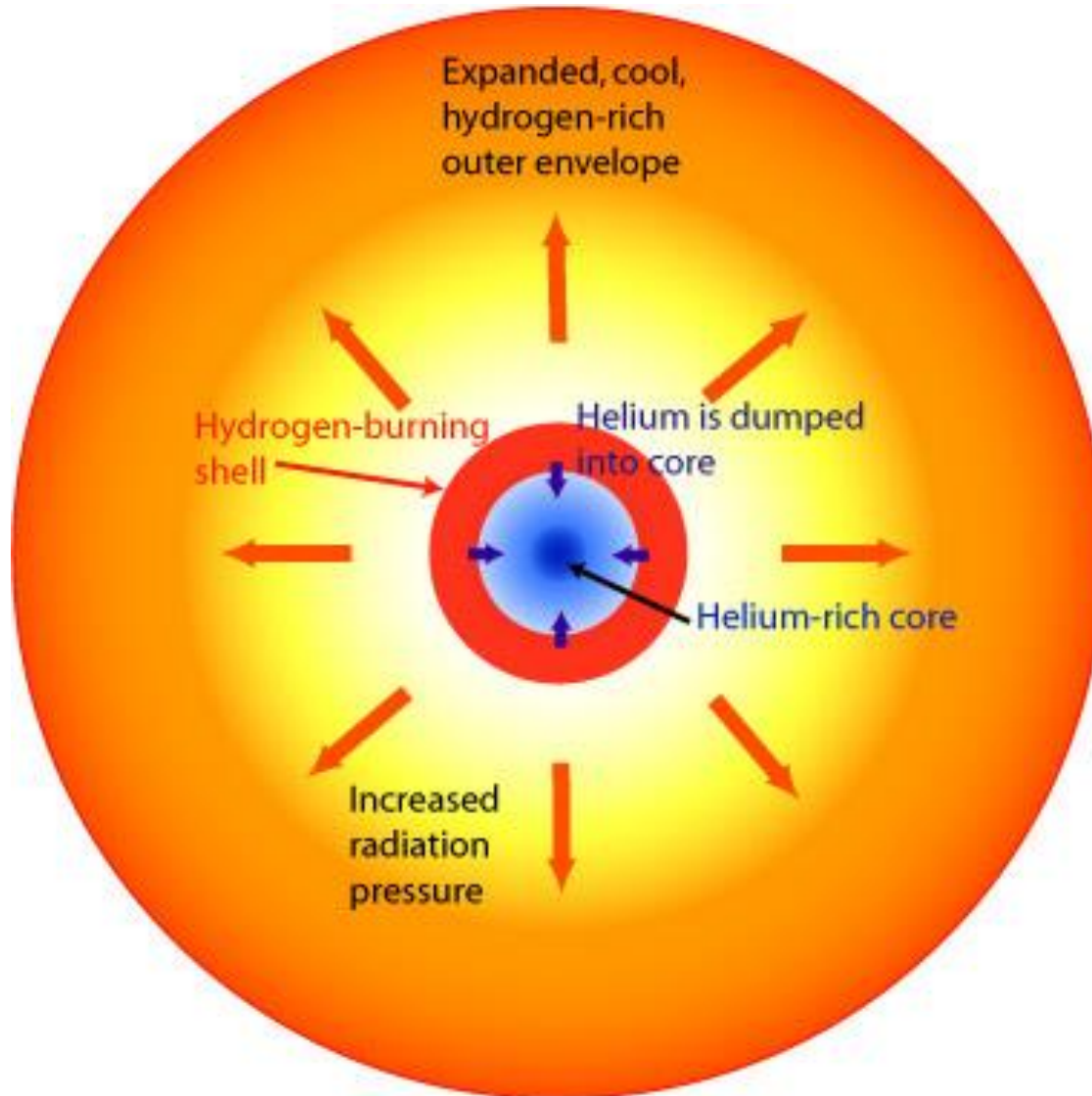
# First dredge-up



Composition depending on mass  
initial metallicity, and used  
theoretical model

$1 M_{\odot}$	Initial	After FDU
Y	0.280	0.3025
C/Z	0.1733	0.1530
N/Z	0.0531	0.0771
O/Z	0.4823	0.4822
${}^3\text{He}$	8.4E-5	0.001368
${}^{12}\text{C}/{}^{13}\text{C}$	90.0	29.8
${}^{16}\text{O}/{}^{17}\text{O}$	2660	2610
${}^{16}\text{O}/{}^{18}\text{O}$	500.1	526.9

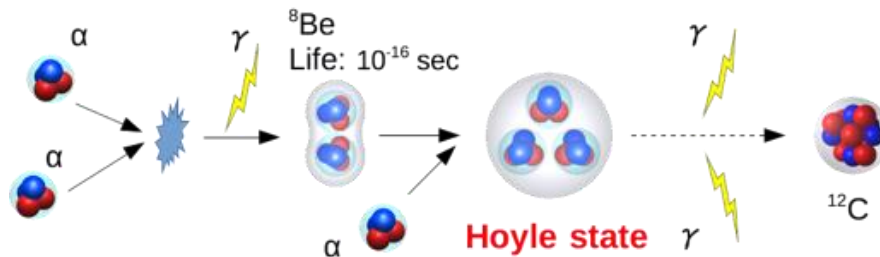
# Red Giant Phase



Hydrogen Shell Burning on the Red Giant Branch

# Red Giant Phase

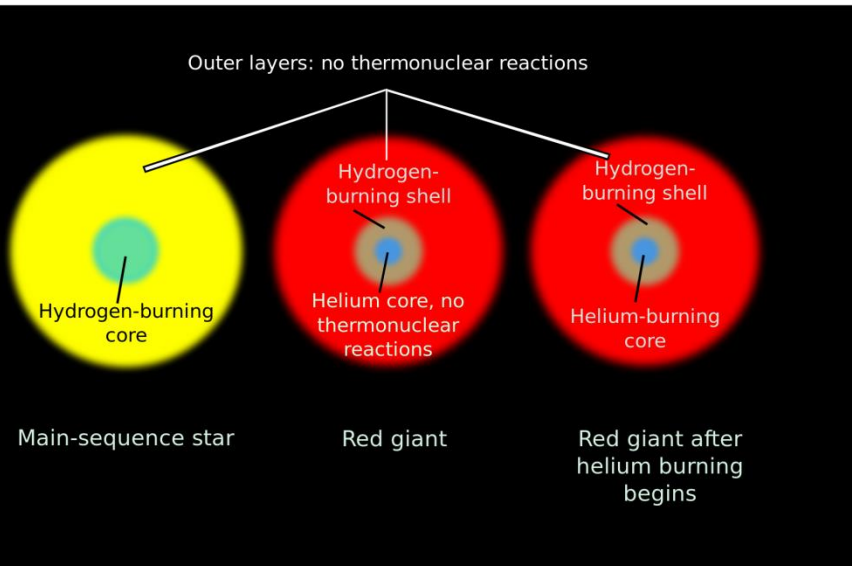
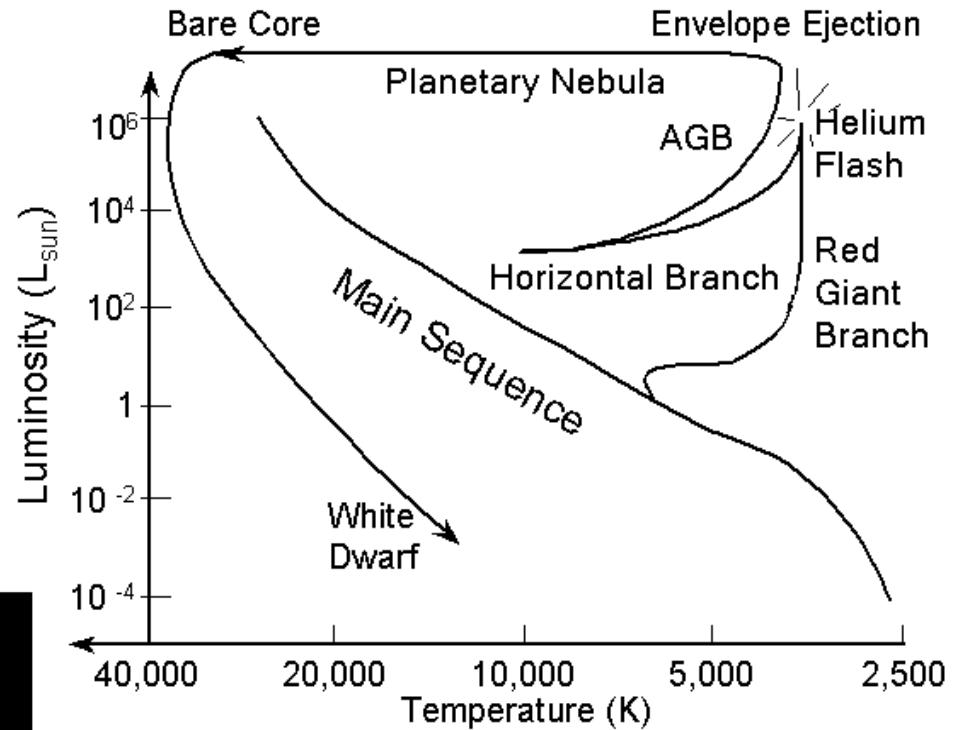
- With time, more helium “ash” adds into the core
- Core contracts more and becomes even hotter
- When the central temperature reaches  $10^8$  K, helium fusion ignites (***Helium Flash***) inside the core
- Helium fusion process (triple alpha process), converts helium to carbon





# Helium Flash

Helium Flash takes only seconds



# Helium Flash

- In low-mass star, the ***compressed core*** is in an ***electron-degeneracy state***
- ***Electron-degeneracy***: the ***electrons*** are so closely packed that they ***can not be further compressed***, due to the ***Pauli exclusion principle***
- ***Pauli exclusion principle***: two particles can not occupy the same quantum state
- The ***core*** becomes supported by ***degenerate-electron pressure***, i.e ***independent of temperature***
- As ***helium fusion ignites*** in the core, ***temperature rises exponentially***, but ***pressure does not rise***

# Helium Flash

- ***Helium fusion rate rises exponentially*** => Helium Flash
- During the time of Helium Flash, the ***core*** is ***extremely bright*** ( $10^{12} L_{\odot}$ )
- At certain core temperature ***no longer electron-degeneracy state*** => core becomes an ***ideal gas***
- The ***core expands***, and ***cools***, terminating the Helium Flash
- The ***core*** becomes a ***steady state*** of ***helium fusion***
- A red giant enters into the stage of “***Horizontal Branch***” in the HRD

# Horizontal Branch

- Red Giant becomes *less luminous* and *smaller*, but *hotter*
- ***Steady State*** of helium fusion in core: *no contraction* and *no temperature rising*
- The ***core*** acts like an ***ideal gas***:
  1. Temperature increases, core expands
  2. Core expands, temperature decreases
- Shell hydrogen fusion rate drops => lower luminosity
- Red Giant shrinks because of less energy output
- It becomes hotter at surface as it compresses